A concept for the global assessment of tomographic resolution and uncertainty

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17 May 2024

SUMMARY

A major challenge in seismic tomography consists in quantifying and representing model resolution and uncertainty, particularly at global scales. This information is crucial for interpretations of tomographic images and their technical application in geodynamics. However, due to large computational costs, there have been only few attempts so far to coherently analyse the spatially varying resolving power for a complete set of model parameters. Here, we present a concept for an effective evaluation and global representation of the 3-D resolution information contained in a full set of averaging kernels. In our case, these kernels are constructed using the 'Subtractive Optimally Localized Averages' (SOLA) method, a variant of classic Backus-Gilbert inversion suitable for global tomography. Our assessment strategy incorporates the following steps: 1) a 3-D Gaussian function is fitted to each averaging kernel to measure resolution lengths in different directions; 2) we define a classification scheme for the quality of the averaging kernels based on their focus with respect to the estimated 3-D Gaussian, allowing us to reliably identify whether the inferred resolution lengths are robust. This strategy is not restricted to SOLA inversions, but can, for example, be applied in all cases where point-spread functions are computed in other tomographic frameworks.

Together with model uncertainty estimates that are derived from error propagation in the SOLA method, our concept reveals at which locations, resolution lengths and interpretations of model values are actually meaningful. We finally illustrate how the complete information from our analysis can be used to calibrate the SOLA inversion parameters —locally tunable target resolution kernels and trade-off parameters— without the need for visual inspection of the individual resulting averaging kernels. Instead, our global representations provide a tool for designing tomographic models with specific resolution-uncertainty properties that are useful in geodynamic applications, especially for linking seismic inversions to models of mantle flow.

Key words: Seismic tomography - Inverse theory - Body waves - Structure of the Earth

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1 INTRODUCTION

Global seismic tomography is the primary technique for re-2 vealing the physical structure of the deep Earth. The first 3 tomographic models of the Earth's mantle, developed more 4 than four decades ago, have mainly been concerned with 5 mapping seismic heterogeneity at spherical harmonic de-6 grees of 6-8; that is, at wavelengths of thousands of kilome-7 tres (Sengupta & Toksöz 1976; Aki et al. 1977; Dziewonski 8 et al. 1977). Over the years, the resolution of global tomo-9 graphic images has steadily been improving by the general 10

increase in data coverage, by exploitation of datasets with complementary sensitivity, as well as through advanced forward and inverse modelling techniques (e.g. Ritsema et al. 2011; Schaeffer & Lebedev 2013; Zaroli et al. 2015; French 15 & Romanowicz 2015; Koelemeijer et al. 2016; Fichtner et al. 2018; Lu et al. 2019; Hosseini et al. 2020; Lei et al. 2020).

Still, in many regions there is only little consensus on the seismic heterogeneity at shorter length scales of ~ 300 -500 km and less. Not only the exact geographic distribution, but in particular the magnitudes of heterogeneity, are diffi-

cult to constrain with tomographic methods. For example, 21 regularization with damping and smoothing constraints is 22 typically needed to counteract the ill-posed nature of the 23 problem, but this inevitably biases the recovered model am-24 plitudes (e.g. Ritsema et al. 2007; Nolet 2008; Schuberth 25 et al. 2009a). Further intricacies arise from complex non-26 linear wavefield effects and trade-offs between physical pa-27 rameters (e.g Hung et al. 2001; Favier et al. 2004; Zhang & 90 28 Shen 2008; Mercerat & Nolet 2012; Schuberth et al. 2015; 29 Koroni et al. 2022). The continuing desire in tomographic 92 30 studies to increase resolution beyond the current limits is 93 31 not an end in itself, but for global applications motivated 94 32 by the estimated thickness of the thermal boundary layers 33 of the mantle and the associated expected size of slabs and 34 plumes. In light of such geodynamic considerations, an ac-35 curate retrieval of heterogeneity at spatial scales of ~ 100 98 36 km and less is crucial for subsequent quantitative infer-37 ences in studies of the lower mantle (e.g. Schuberth et al. 100 38 2009b; Koelemeijer et al. 2018; Choblet et al. 2023; Richards 101 39 et al. 2023), reconstructed time evolution of mantle flow (e.g. 102 40 Bunge et al. 2003; Spasojevic et al. 2009; Shephard et al. 103 41 2010; Horbach et al. 2014; Colli et al. 2018; Ghelichkhan 104 42 et al. 2021) and surface dynamic topography (e.g. Davies 105 43 et al. 2019, 2023). In addition to the dynamically inherent 106 44 size of thermal anomalies, variations in mineral phase assem-107 45 46 blage and chemical composition likely occur on even shorter 108 47 scales (e.g. Stixrude & Lithgow-Bertelloni 2007; Papanag-109 nou et al. 2022). 48 110

49 Despite the great progress in global seismic tomogra-111 50 phy, relatively few studies addressed explicitly the problem 112 51 of quantifying the spatially variable resolving power of a 113 given inversion (e.g. Boschi 2003; Ritsema et al. 2004; Sol-52 114 dati et al. 2006; Ritsema et al. 2011; Koelemeijer et al. 2016; 115 53 Simmons et al. 2019). Those studies have in common that 116 54 their tomographic systems are based on a linearization of 117 55 the (weakly) non-linear problem, such that the quantifica-118 56 tion of resolution is straightforward from a theoretical point 119 57 of view. This requires the computation of the resolution ma-120 58 trix, which allows for a complete characterization of the un-121 59 derlying effects of imperfect data coverage and regulariza-60 122 tion. Moreover, the linear nature of the solution lends itself 61 123 to practical applications where limited resolution is a critical 124 62 aspect of quantitative model interpretation. An example for 125 63 this is tomographic filtering of geodynamic Earth models, 126 64 which is a necessary step to obtain fair and consistent com-65 127 parisons between these independent theoretical predictions 66 128 of present-day seismic heterogeneity and the tomographi-129 67 cally imaged structures (e.g. Mégnin et al. 1997; Ritsema 130 68 et al. 2007; Schuberth et al. 2009a; Nerlich et al. 2016; Sim-69 131 mons et al. 2019; Freissler et al. 2020). Together with the 70 132 posterior covariance matrix, which includes the variances 71 133 and correlations of model parameters, the non-uniqueness 72 134 and quality of the tomographic solution can be fully ap-73 135 praised (e.g. Nolet 2008; Simmons et al. 2019). 74 136

For non-linear systems, in contrast, a complete formal 75 137 quantification of resolution and uncertainty is often not vi-138 76 able in practice. In full waveform inversions based on nu-139 77 merical wavefield simulations and adjoint techniques (e.g. 140 78 Igel et al. 1996; Pratt 1999; Fichtner et al. 2009; Tape et al. 141 79 2009; Colli et al. 2013; Krischer et al. 2018; Ma et al. 2022; 142 80 Rodgers et al. 2022), there is still a gap between the relative 143 81 wealth of information in the data and the available tools to 144 82

assess the general non-uniqueness, especially in global models. One possibility to approach this problem is Bayesian inference (Tarantola 2005) in order to elegantly deal with the non-linearity. However, probabilistic approaches that rely on sheer random sampling of the posterior probability density function are out of reach for global scale applications due to the high computational costs of repeatedly solving the forward problem in such analyses. Several other strategies have thus been proposed to address this issue, which mostly involve an approximation of the Hessian matrix for the misfit function in the vicinity of the 'optimal' model. This is motivated by the insight that the Hessian matrix may be exploited in a local sense as the inverse of the posterior model covariance (e.g. Tarantola 2005; Bui-Thanh et al. 2013; Liu et al. 2021). It allows for some practical inferences, such as in extremal bounds analysis (Fichtner 2010), or for an efficient exploration of the model nullspace (e.g. Deal & Nolet 1996; Liu & Peter 2020; Fichtner et al. 2021). Hessian-vector products may also be used to compute local point-spread functions (e.g. Fichtner & Trampert 2011; Fichtner et al. 2013) that are equivalent to the columns of the resolution matrix in a linear framework. However, except for pointspread function tests for a few individual locations in the full-waveform models GLAD-M15 (Bozdağ et al. 2016) and GLAD-M25 (Lei et al. 2020), these sophisticated methods have found only limited usage in global scale applications so far.

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It must be noted that even in the linear case, computing formal resolution and uncertainty is a formidable challenge (e.g. Rawlinson et al. 2014). Stochastic techniques may yield specific characteristics of the resolution matrix, such as depth-dependent average resolution lengths (Trampert et al. 2013) or the main diagonal elements (MacCarthy et al. 2011). The diagonal entries give an indication of the resolvability at the parameter location of interest, while resolution lengths characterize the impact range of off-diagonal entries representing inter-parameter trade-offs. More detailed information can be extracted, for example by a statistical resolution matrix (An 2012), or a stochastic estimation of pointspread function parameters, which can in turn be applied to both linear and non-linear problems (Fichtner & Leeuwen 2015). It is also possible in large-scale problems to use direct approaches for computing the resolution matrix that involve efficient numerical strategies (Boschi 2003; Soldati et al. 2006; Bogiatzis et al. 2016). Alternatively, practical approximations can be made to estimate both the resolution matrix and the posterior covariance (e.g. Nolet et al. 1999; Simmons et al. 2019).

A straightforward method to compute directly the resolution as well as uncertainty can be found in the seminal work by Backus & Gilbert (1967, 1968, 1970). In Backus–Gilbert theory, the estimates of individual model parameters can be interpreted as localized spatial averages around a given target location. In contrast to the more commonly used linear methods in tomography, which often employ Tikhonov regularization for norm damping, no a priori constraints on model values need to be prescribed that may bias model amplitudes. Instead, a certain control can be exerted on the trade-off between a favourable spatial structure of averaging kernels (that determine resolution) and the amount of data errors propagating into the averages as model uncertainties (Backus & Gilbert 1970).

A variant of the Backus-Gilbert method, called Sub- 207 145 tractive Optimally Localized Averages (SOLA), was intro-208 146 duced to global seismic tomography by Zaroli (2016). Orig-209 147 inally formulated and termed SOLA by Pijpers & Thomp-210 148 son (1992, 1994) for 1-D inversions in helioseismology, the 149 211 method may have been discovered independently by several 150 212 authors in different contexts (e.g. Oldenburg 1981; Louis 213 151 & Maass 1990). In geophysics, it was further adapted for 214 152 solving discrete and continuous 2-D and 3-D large-scale to- 215 153 mographic problems (Zaroli et al. 2017; Zaroli 2019) and 216 154 has since been applied to surface wave tomography (Latal- 217 155 lerie et al. 2022; Amiri et al. 2023), normal modes (Restelli 218 156 et al. 2024), and modelling the radial magnetic field at the 219 157 core-mantle boundary (Hammer & Finlay 2019). The great 220 158 advantage of the SOLA method compared to the classic 221 159 Backus–Gilbert formulation arises from the implementation 222 160 of target kernels with prescribed finite size, which specify 161 the volume around the specific parameter location in which 162 the inversion shall provide the spatial average of the model 163 223 values. The target kernels make it possible to provide a pri-164 224 ori information on the expected local resolution length scales 165 (i.e. they allow for potentially exerting a rather direct con-166 trol on the final resolution), while the so-called trade-off pa-167 rameter moderates error propagation. Moreover, the SOLA 168 method enables perfectly parallel computations of the model 169 170 values as well as of the averaging kernels and propagated 171 uncertainty. So far, however, there is no definite method or recipe for the automatic selection of the SOLA inversion 172 parameters, namely the size of the individual target ker-173 nels and the particular choice for the trade-off parameters, 174 throughout a complete model. In global seismic tomogra-175 phy, such a strategy would be particularly helpful due to 176 the highly inhomogeneous data coverage, leading to locally 177 different quality of the tomographic solution. The remaining 178 issue in that regard is the lack of tools for assessing the entire 179 set of averaging kernels in a 3-D setting. Furthermore, even 180 if one can compute with SOLA, or any other tomographic 181 method, a complete set of averaging kernels (or point-spread 182 functions), one will never be able to visually calibrate and 183 analyse each one individually. In other words, it is still a 184 225 challenge in itself to effectively represent and communicate 185 226 the resolution information (Trampert 1998). 186

227 The objective of this paper is therefore twofold: First, 187 228 we want to explore for a previously employed tomographic 188 dataset, how different inversion parameter choices in the 229 189 SOLA method applied to global S-wave tomography lead 230 190 to different local resolving power and model uncertainties. 191 231 To this end, we systematically test several combinations of 192 target kernels and trade-off parameters spanning the range 193 from low-resolution to high-resolution inversions, each with ²³² 194 varying degrees of resulting model uncertainty. Second, in 195 order to effectively analyse the results from different inver-²³⁴ 196 sion parameter combinations, we develop a combined anal-235 197 ysis of the resolution length scales in 3-D together with a $_{\rm 236}$ 198 specific test of the adequateness of the method for estimat- ²³⁷ 199 ing these lengths. This allows us to represent the practically ²³⁸ 200 relevant information on resolution in the SOLA averaging ²³⁹ 201 kernels on a global scale, which can then be inspected along-²⁴⁰ 202 side the uncertainty propagating into the model. 203

We start with a brief review of the SOLA method and describe the general tomographic system that we employ in Section 2 and Appendix A. Computational aspects regard-242

ing the efficient solution of the linear SOLA system are described in Appendix B. Section 3 provides examples of typical SOLA averaging kernels for different inversion choices and motivates the development of a strategy for estimating resolution lengths with a Gaussian approximation in Section 4. Important for this analysis will be to test this Gaussian approximation, for which we introduce the concept of 'focus' that allows us to define different quality categories for the averaging kernels. In Section 5, we provide global maps of tomographic resolution lengths in specific but globally coherent directions, estimated in a robust manner using the combined concepts of resolution and focus, together with the propagated uncertainty. Finally, we discuss the impact of the different possible choices of inversion parameters in the SOLA method in light of possible optimal design towards practical applications.

2 TOMOGRAPHIC METHODOLOGY

2.1 Review of the SOLA Backus–Gilbert method

The main insight of Backus–Gilbert theory is relatively straightforward: given a finite amount of data, one can generally not retrieve exact point estimates of the Earth model parameters $m(\mathbf{r})$ of interest. Nevertheless, it is often possible to infer at a model target location $\mathbf{r}^{(k)}$ a unique weighted average $\hat{m}^{(k)}$, such that

$$\hat{m}^{(k)} = \int_{V} A^{(k)}(\mathbf{r}) \, m(\mathbf{r}) \, d^{3}\mathbf{r} \,, \qquad (1)$$

where $A^{(k)}(\mathbf{r})$ is the averaging or resolving kernel (Backus & Gilbert 1968, 1970). Classically, the objective is to obtain an optimally localized averaging kernel that approximates a delta peak at $\mathbf{r}^{(k)}$, constructed from a linear combination of N data(-sensitivity) kernels $K_i(\mathbf{r})$. In the linear case, the data kernels $K_i(\mathbf{r})$ relate model parameters $m(\mathbf{r})$ to the measured data d_i in the form of

$$d_{i} = \int_{V} K_{i}(\mathbf{r}) m(\mathbf{r}) d^{3}\mathbf{r} + n_{i}, \quad i = 1, ..., N, \qquad (2)$$

where the data d_i include an error (or noise) component n_i that is assumed here to be independent and normally distributed with zero mean and variance $\sigma_{d_i}^2$. To compute global sets of averaging kernels $A^{(k)}(\mathbf{r})$, we employ the SOLA method (Zaroli 2016) that solves the following optimization problem:

$$\min_{\mathbf{x}^{(k)}} \int_{V} \left(A^{(k)}(\mathbf{r}) - T^{(k)}(\mathbf{r}) \right)^{2} d^{3}\mathbf{r} + \eta^{2} \sigma_{\hat{m}^{(k)}}^{2}$$
(3)

subject to
$$\int_{V} A^{(k)}(\mathbf{r}) d^{3}\mathbf{r} = 1$$
, (4)

where $T^{(k)}(\mathbf{r})$ is a target (resolution) kernel, η the trade-off parameter and $\sigma^2_{\hat{m}^{(k)}}$ the model variance from error propagation. The solution of eq. (3)+(4) yields a set of coefficients $\mathbf{x}^{(k)}$ that can be interpreted as (the k-th row of) a generalized inverse operator and determines the estimated average $\hat{m}^{(k)}$, the averaging kernel $A^{(k)}(\mathbf{r})$ and the model uncertainty from error propagation $\sigma_{\hat{m}^{(k)}}$ (plus covariance if desired):

$$\mathbf{x}^{(k)} \Longrightarrow \begin{cases} \sum_{i=1}^{N} x_i^{(k)} d_i & \longrightarrow \hat{m}^{(k)} \\ \sum_{i=1}^{N} x_i^{(k)} K_i(\mathbf{r}) & \longrightarrow & A^{(k)}(\mathbf{r}) \\ (\sum_{i=1}^{N} (x_i^{(k)})^2 (\sigma_{d_i})^2)^{1/2} & \longrightarrow & \sigma_{\hat{m}_k} \,. \end{cases}$$
(5)



Figure 1. Left: Cross-section of a target kernel with horizontal and vertical half widths $w_{\rm H/V} = 600/300$ km. The cyan ellipse marks the contour line at half maximum. Right: Cross-section for an averaging kernel at the same target location. To get an idea of the fit to the target kernel, we also plot the target ellipse at half maximum. 298

The unimodular constraint in eq. (4), which is also implied ²⁹⁹ 243 by the classical Backus–Gilbert theory, ensures that $\hat{m}^{(k)}$ 300 244 represents a physical volumetric average. In the hypothetical ₃₀₁ 245 presence of a constant model value around the target loca-302 246 247 tion $\mathbf{r}^{(k)}$, the estimated model amplitude would thus not be 303 artificially scaled. Most important, the introduction of a tar-304 248 get kernel $T^{(k)}(\mathbf{r})$ in the SOLA method means that, instead 249 305 250 of an ideal delta peak, as in the original Backus–Gilbert for-306 mulation, a practically more relevant spatial function is un-251 307 derlying the construction of the averaging kernels. Different $_{\ 308}$ 252 choices of $T^{(k)}(\mathbf{r})$ then allow us to use information on the ex-253 309 pected local resolution by varying the shape and weighting ₃₁₀ 254 present in the target kernels. At the same time, the trade-255 311 off parameter η ensures that the fit to the target kernel and 312 256 the variance of the propagated errors in the inferred aver-257 ages can be controlled. Both $T^{(k)}(\mathbf{r})$ and η may be selected 314 258 with a subjective preference, however with the benefit that 315 259 the inversion results include complete information on reso-260 316 lution and uncertainty. Note also that each choice of η and 261 317 $T^{(k)}(\mathbf{r})$ has a specific impact on all local model properties 262 318 (for details on the SOLA inversions employed here, includ-263 319 ing computational aspects to solve the system efficiently for 320 264 all model parameters, see Appendix A and B). 265 321

266 2.2 Resolution length and choice of target kernels

327 To obtain robust resolution information from the SOLA 267 averaging kernels, Pijpers & Thompson (1994) calibrated 268 a practical threshold value based on visual inspection of 269 328 their target fit. This way, they were able to distinguish well-270 localized from non-localized 1-D averaging kernels and could 271 329 use the prescribed target kernel peak width to infer resolu-330 272 tion lengths. This is however not directly applicable in our 331 273 case. In 3-D tomography, the fit of $A^{(k)}$ to the target reso-332 274 lution might be good in a specific direction and poor in any 333 275 other one. Therefore, no simple choice of target kernel $T^{(\vec{k\,)}}$ 334 276 and trade-off parameter η is reducing equally well the kernel 335 277 difference in all regions. Instead, additional information on 336 278 possible shifts and varying length scales in different direc- 337 279 tions is required to assess the averaging kernels (see Section 338 280 4). However, with the choice of a specific target $T^{(k)}$, one 339 281 can still promote a desired shape and size of the $A^{(k)}$ prior $_{340}$ 282 to the inversion. 283 341

For this purpose, an intuitive parametrization for the target kernel $T^{(k)}$ is given by 3-D Gaussian functions that have been used previously in seismology for analysing tomographic resolution (see e.g. Fichtner & Trampert 2011; An 2012). The kernels are centred at the target locations $\mathbf{r}^{(k)}$, which correspond to the points in a tomographic grid (see Section 2.3) By using a local Cartesian frame with origin at $\mathbf{r}'^{(k)}$ (the prime indicates the change of basis), we can align the principal axes of the Gaussian function along the horizontal and vertical directions; i.e. they are oriented along tangents in the latitudinal and longitudinal direction and along the radius in the vertical direction, with respect to $\mathbf{r}'^{(k)}$. Our 3-D Gaussian target kernels are then given by

$$T^{(k)}(x',y',z') = \frac{a^3}{\sqrt{(2\pi)^3} \cdot w_{x'} \cdot w_{y'} \cdot w_{z'}} \times \exp\left[-\frac{a^2}{2} \left(\frac{x'^2}{w_{x'}^2} + \frac{y'^2}{w_{y'}^2} + \frac{y'^2}{w_{y'}^2}\right)\right]$$

where $w_{x',y',z'}$ are the half widths at half maximum and x', y' and z' refer to the axes in the local basis. Examples for a target kernel $T^{(k)}$ and a resulting averaging kernel $A^{(k)}$ are visualized in Fig. 1. We specify the $w_{x',y'}$ in horizontal $(w_{\rm H})$ and $w_{z'}$ in vertical direction $(w_{\rm V})$ of $T^{(k)}$ as target resolution lengths. As a remark, their relation to the standard deviation is $w = a \cdot \sigma$, where $a = \sqrt{2 \ln(2)} \approx 1.17$. Using w is particularly useful because it allows one to relate the kernel width to its maximum at the peak, which facilitates comparisons of the volumetric change of various kernels with different peak amplitudes. We directly evaluate eq. (8) on the discrete tomographic grid (see Section 2.3). In contrast to Zaroli (2016), the $T^{(k)}$ are here not strictly normalized, i.e. we do not enforce $\int_V T^{(k)}(\mathbf{r}) d^3 \mathbf{r} = 1$. This is done deliberately in order to preserve the Gaussian shape of $T^{(k)}$ within the finite volume V of the grid.

For the SOLA inversions in this study, we take 3 variations of $T^{(k)}$ using Gaussian functions of progressively larger horizontal and vertical target resolution lengths ($w_{\rm H/V} = 300/200, 600/300, 900/400$ km). Along with this, we test 3 different values for the trade-off parameter η ($\eta_1 = 5, \eta_2 = 10, \eta_3 = 30$; see also Table 1). Increasing values for η generally promote smaller model uncertainty $\sigma_{\hat{m}^{(k)}}$ while deteriorating the fit of the averaging kernel $A^{(k)}$ to $T^{(k)}$. The range of values of η was chosen empirically from a few SOLA inversions for all $T^{(k)}$ in order to cover a range of tomographically relevant levels for $\sigma_{\hat{m}^{(k)}}$, resolution length scales and target fits.

2.3 Tomographic grid and dataset

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Since we are interested in quantifying the impact of different inversion parameter choices, we choose a global grid that can well represent the shape of averaging kernels down to all target length scales; that is, it covers at least the smallest target half widths $w_{\rm H/V} = 300/200$ km used here. In general, we follow the parametrization strategy of Zaroli (2016), where grid nodes are the upper vertices of triangular prisms based on a spherical Delaunay triangulation for several distinct depth layers across the entire mantle. For details, the reader is referred to Zaroli (2010, 2016). In radial direction, we take the 18 depth layers from SOLA-Z16 (Zaroli 2016) that are between 100-200 km thick, whereas in lateral direction, we use approximately equidistant spherical Fibonacci



Figure 2. Left: Fibonacci grid on a sphere for the 530-660 km depth layer using the algorithm by Swinbank & Purser (2006). The nodes are the upper vertices of spherical triangular prisms that are constructed by Delaunay triangulation. Right: Histogram of internode distances for all 18 depth layers. The total amount of unique lateral node connections (i.e. the number of triangle edges) covered in the histogram is 447,567. The average distance is \sim 224 km. 382

384 grids following Swinbank & Purser (2006). In Fig. 2 the grid 342 385 nodes of the layer at 530-660 km depth are shown as an 343 example. To create a suitable realization of the Fibonacci ³⁸⁶ 344 grid, we empirically determine a specific amount of points ³⁸⁷ 345 for each layer such that the maximum distance of neigh-346 bours from the spherical Delaunay triangulation is less than ³⁸⁹ 347 390 300 km. The distribution of the node distances across all 348 layers is shown in a histogram in Fig. 2. Minimum and av- $^{\rm 391}$ 349 erage neighbour distances are about 176 km and 224 km, $^{\rm 392}$ 350 393 respectively. In total, the grid includes 153,323 grid nodes. 351

394 As tomographic dataset, we use the source-receiver con-352 395 figurations from the SOLA-Z16 model (Zaroli 2016) that in-353 clude 79,765 cross-correlation traveltime measurements for S 354 and SS seismic phases at 22.5 s central period (Zaroli 2010). 397 355 398 Sensitivities to shear-wave velocity perturbations are calcu-356 399 lated using paraxial finite-frequency kernels (Dahlen et al. 357 400 2000). The coverage of the dataset is particularly suitable 358 for investigating the velocity structure at depths between ⁴⁰¹ 359 402 400 and 1710 km (Zaroli 2016). However, there are mainly 360 403 two relevant regions that have different characteristic be-361 404 haviour for resolution. At depths of $\approx 400-810$ km the finite-362 frequency kernels for teleseismic S-waves have not bottomed 405 363 out yet, thus finding suitable linear combinations of the data 364 to enhance resolution locally can be more difficult there. In $^{\rm 407}$ 365 408 contrast, at depths below ≈ 810 km, larger horizontal vol-366 umes are covered by the finite-frequency kernels, leading to ⁴⁰⁹ 367 a higher chance for crossing volumes of sensitivity. To effec- 410 368 tively investigate the resulting averaging kernels, we there-369 fore focus on two particular depth layers in this study (at $^{\rm 412}$ 370 530–660 km, midpoint at 595 km and 1110– 1310 km, mid- $^{\scriptscriptstyle 413}$ 371 414 point at 1210 km depth) representing each situation. 372 415

373 3 SOLA AVERAGING KERNELS

To first get a rough appreciation of the behaviour of the SOLA averaging kernels, we visually inspect some examples before we proceed with the global analysis of all kernels in Section 4. To this end, we show kernel cross-sections for the **Table 1.** SOLA inversion parameters used in this study. To facilitate discussion, a high resolution(HR) and a low resolution (LR) setup are introduced. Horizontal and vertical target resolutions $w_{\rm H/V}$ correspond to the half widths at half maximum of the Gaussian target kernels (see eq. 6).

Target resolution $w_{\rm H/V}$ (horizontal/vertical)	Trade-off parameter η
$300/200 {\rm ~km}$	$\eta_1 = 5$
600/300 km	$\eta_2 = 10$
$900/400 \mathrm{~km}$	$\eta_3 = 30$

High resolution (HR) setup : 300/200 km, η_1 Low resolution (LR) setup: 900/400 km, η_3

different combinations (see Table 1) of trade-off parameters η and target kernels $T^{(k)}$ that we introduced in Section 2.2.

3.1 Influence of data coverage

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A compilation of different averaging kernels is shown in Fig. 3 for two end-member combinations of $T^{(k)}$ and η . They represent a high-resolution (HR) inversion with η_1 (= 5) and a target kernel size of $w_{\rm H/V} = 300/200$ km (Figs 3a and d), and the contrary case of a low-resolution (LR) inversion using η_3 (= 30) and a target size of $w_{\rm H/V} = 900/400$ km (Figs 3b and c). Dashed ellipses defined by horizontal and vertical semi-axes $w_{\rm H}$ and $w_{\rm V}$ indicate the size of $T^{(k)}$ for the particular inversion (as shown in Fig. 1). Data coverage is particularly good in the Northern Hemisphere (top row in Fig. 3), where one can see that both HR and LR inversion setups lead to averaging kernels that are overall well localized and fit the shape of the target kernels. In Fig. 3(b) using the LR setup for an averaging kernel centred below North America, one can still observe stronger positive amplitudes south of Hawaii outside the broader target region. Poor data coverage in the Southern Hemisphere (bottom row in Fig. 3) typically leads to increased power in side lobes, reflecting the difficulty of fitting the target kernel with a locally incomplete dataset. For example, the averaging kernel in Fig. 3(c) for the LR setup is hardly centred and apparently dominated by a subset of unidirectional data kernels reaching to the surface between Kerguelen Islands and Australia. On the other hand, the $A^{(k)}$ in Fig. 3(d) combines poor data coverage with the HR setup, leading again to more pronounced side lobes, but also to a good fit to the target for the bulk of the averaging sensitivity. Fig. 3 already suggests that the target kernel has a strong control on the resolution properties of $A^{(k)}$ (as expected), and that poor data coverage, although making the averaging kernels more prone to oscillatory behaviour, is not necessarily preventing one from fitting the target resolution. Although being barely visible here, negative values in the averaging kernels do exist, but their amplitudes are mainly located outside the target region and are generally small. A more quantitative analysis taking this into account is presented in the classification scheme that we develop in Section 4.2.

3.2 Variable target kernel and trade-off parameter



Figure 3. Lateral and vertical cross-sections for four different averaging kernels in regions of good and poor data coverage in the Northern and Southern Hemisphere, respectively. Left and right panels show target locations at 595 km and 1210 km depth, respectively. a) and d): 'High-resolution' setup (HR), $A^{(k)}$ for a trade-off parameter of $\eta_1 = 5$ and target kernel $T^{(k)}$ with horizontal/vertical extent of $w_{H/V}$ = 300/200 km. b) and c): 'Low-resolution setup (LR)', $A^{(k)}$ for $\eta_3 = 30$ and $T^{(k)}$ with $w_{H/V} = 900/400$ km. The low-resolution setup leads to a preference of lower uncertainty $\sigma_{\hat{m}^{(k)}}$ by using a larger trade-off parameter η_3 at the cost of a worse fit to the target resolution. The latter is marked by dashed ellipses (drawn at the half widths at half maximum $w_{H/V}$ of each $T^{(k)}$). Values of the averaging kernels are normalized with their respective maximum. Because our tomographic grid is fine enough for the target length scales, it follows that for an almost ideal fit, the darkest red values at the cut-off of the colour scale would all be within the target ellipses. Black contour lines in the vertical cross-sections further indicate iso-surfaces at 1/2, 1/4, 1/8, 1/16 and 1/32 of the peak value.

While Fig. 3 visualizes the consequences of variable data 438 421 coverage, we can also systematically exploit the capabilities 439 422 of the SOLA method and investigate how different choices 440 423 of target kernel and trade-off parameter affect $A^{(k)}$ and the 441 424 propagated uncertainty $\sigma_{\hat{m}^{(k)}}.$ Fig. 4 features such a set $_{\rm 442}$ 425 of alternatively possible averaging kernels, corresponding to $_{443}$ 426 the target location and the low-resolution example of Fig. $_{\rm 444}$ 427 3(c) in the Southern Hemisphere. Here, we focus on vertical $_{445}$ 428 cross-sections in west-east direction; complementary figures 446 429 for all cross-sections and kernels can be found in the online $_{447}$ 430 supplementary material. The different averaging kernels are $_{448}$ 431 organized in a matrix layout, going from smaller to larger 449 432 433 target kernel sizes from top to bottom and increasing values 450 434 for the trade-off parameter η from left to right. 451

An increase in η leads to an increase in model uncertainty $\sigma_{\hat{m}^{(k)}}$, which is indicated in each cross-section in Fig. 4, ranging from 0.47 % for the HR setup, down to

0.03 % dln(v_S) for the LR setup. Likewise, at constant η and growing target kernel sizes, model uncertainties also decrease. None of the averaging kernels exactly attains the maximum amplitude of their underlying $T^{(k)}$. At a target size of $w_{\rm H/V} = 300/200$ km, for example, the maxima of the corresponding $A^{(k)}$ are off by more than a factor of 2. For larger target lengths, this difference is less severe. Overall, an unsatisfactory visual fit of the averaging kernels to their respective target kernels is observed, which can be ascribed here to a single finite-frequency kernel that seems to dominate the averaging. A different situation is shown in Fig. 5. At this location, the data coverage is excellent, which leads to a suite of averaging kernels that are Gaussian-shaped for all parameter combinations tested. Peak amplitudes of the averaging kernels for the lowest value of the trade-off parameter are close to their respective target kernel value and only moderately drop as η is increased. Even at lower



Resolution in global seismic tomography 7

Figure 4. Influence of increasing the trade-off parameter and target kernel size on the averaging kernels for a region of relatively poor data coverage in the southern Indian Ocean, east of the Kerguelen Islands at 595 km depth. Left: lateral cross-sections of the target kernel $T^{(k)}$ with horizontal and vertical extent $w_{\rm H}$ and $w_{\rm V}$, respectively, used for the inversion. Right: vertical cross-sections (West-East) of the resulting averaging kernels $A^{(k)}$ for all combinations of η ($\eta_{1,2,3} = 5$, 10, 30) and $T^{(k)}$. Dashed ellipses indicate the size of $T^{(k)}$ with semi-axes of length $w_{\rm H}$ and $w_{\rm V}$. The kernel values are normalized with their respective half maximum to facilitate comparisons. Model uncertainties $\sigma_{\hat{m}^{(k)}}$ are given in % dln(v_S).



Figure 5. Influence of increasing the trade-off parameter and target kernel size for a region of excellent data coverage in the central US at 1210 km depth. Same layout as Fig. 4.



Figure 6. Concept for our resolution analysis and definition of the 'focus' for an averaging kernel $A^{(k)}$ parametrized with a 3-498 D Gaussian. Yellow and blue indicate an uncorrelated and cor-499 related version of the Gaussian, respectively. Filled ellipses at the half widths at half maximum $(w_{x^\prime/y^\prime/z^\prime})$ define estimated 500 resolution lengths, while the larger, unfilled ellipses at the half 501 width at one-eighth maximum define the region used for com-502 puting the focus ξ (see Section 4.2). In contrast to the simpler, 503 uncorrelated Gaussian $\hat{g}^{(k)}$, the Gaussian $\tilde{g}^{(k)}$ with correlation 504 includes rotation with respect to the axes of the local Cartesian 505 frame centred at the target location. This could be used for estimating minimum/maximum resolution lengths $w_{\min/\max}$, which however, would not lie along the same directions for all kernels.

510 model uncertainties $\sigma_{\hat{m}^{(k)}}$, the kernels also mostly retain lo-455 511 456 calization and recover the target shape.

512 It is clear that a visual inspection of the averaging kernels 457 513 might be insufficient to judge on their quality and cannot be 458 514 performed individually for every model parameter. Also, it 459 515 does not provide information on the length scales of resolu-460 tion that need to be quantified for a more thorough analysis. 516 461 517 This raises the question of how one could assess the qual-462 ity and nature of the $A^{(k)}$ in a consistent manner, such that 518 463 519 sensible estimates and comparisons of resolution lengths can 464 at best be made for the entire set of averaging kernels all 465 together. 466 521

QUANTIFYING RESOLUTION 4 467

523 A general problem for inferring meaningful information on 468 resolution in seismic tomography is that one cannot guar-469 524 antee that every individual averaging kernel is localized and 470 525 reasonably Gaussian-shaped for a given selection of inver-526 471 sion parameters. Therefore, we cannot reliably use the pre-527 472 scribed size of the target kernels $T^{(k)}$ to examine resolution 528 473 on a global scale.Instead, resolution lengths should be quan-529 474 tified consistently for all averaging kernels and independent 530 475 of the respective target form of $T^{(k)}$. In addition, it needs 476 531 to be determined whether the resulting length estimates are 477 meaningful in the given context, while bearing in mind the 478 possible complexity of the averaging kernels. A robust quan- 532 479 tification and interpretation of the resolution information 533 480 contained in tomographic inversions thus requires two indi- 534 481

vidual tools: the concept of resolution lengths together with 535 482

a classification scheme for the quality of the averaging ker-483 nels. 484

Gaussian estimates and resolution lengths 4.1

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Owing to our choice of target kernels in the form of 3-D Gaussian functions, it is useful to independently determine also a best-fitting 3-D Gaussian for each averaging kernel in order to quantify resolution lengths. A general 3-D Gaussian parametrization is given by (see also Fichtner & Trampert 2011)

$$\tilde{g}^{(k)}(\mathbf{r}') = \frac{N^* a^3}{\sqrt{(2\pi)^3 \det \mathbf{C}}} \times \exp\left[-\frac{a^2}{2}(\mathbf{r}' - \boldsymbol{\mu}')^{\mathrm{T}} \mathbf{C}^{-1}(\mathbf{r}' - \boldsymbol{\mu}')\right],$$
(7)

where the position \mathbf{r}' and mean vector $\boldsymbol{\mu}'$ are defined in the same local Cartesian coordinate system used for the target kernel $T^{(k)}$ in eq. (6). Any non-zero mean location μ' implies that the centre of mass of the averaging kernel is displaced from the target location (this is called 'distortion' by Fichtner & Trampert 2011). The scaling factor N^* represents the total mass of the Gaussian. We include N^* to facilitate finding suitable resolution length estimates (given by the half widths $w_{x',y',z'}$) that better reflect the averaging volume around the main peak of $A^{(k)}$. The covariance matrix \mathbf{C} describes the spatial extent and rotation/tilting of the Gaussian in the local framework and incorporates the half widths $w_{x',y',z'}$ and parameters for correlation $\rho_{x'y'}$, $\rho_{x'z'}$ and $\rho_{y'z'}$. The complete set of parameters for estimation would thus be given by $\tilde{\mathbf{p}} =$ $(N^*, \mu_{x'}, \mu_{y'}, \mu_{z'}, w_{x'}, w_{y'}, w_{z'}, \rho_{x'y'}, \rho_{x'z'}, \rho_{y'z'})$. By using the full set of correlation parameters, one could therefore also extract from the averaging kernels minimum/maximum resolution lengths in any necessary direction.

For global comparisons of averaging length scales with respect to the uncorrelated target kernels that we have chosen, it might in fact be simpler for interpretation, but not less informative, to focus only on specific directions. To this end, we can neglect the correlations and simplify eq. (7) to get a Gaussian parametrization in the following form:

$$\hat{g}^{(k)}(x',y',z') = \frac{N^* a^3}{\sqrt{(2\pi)^3} \cdot w_{x'} \cdot w_{y'} \cdot w_{z'}} \times \exp\left[-\frac{a^2}{2} \left(\frac{(x'-\mu_{x'})^2}{w_{x'}^2} + \frac{(y'-\mu_{y'})^2}{w_{y'}^2} + \frac{(z'-\mu_{z'})^2}{w_{z'}^2}\right)\right].$$
(8)

The reduced set of parameters for estimation is given by $\mathbf{p} = (N^*, \mu_{x'}, \mu_{y'}, \mu_{z'}, w_{x'}, w_{y'}, w_{z'})$, from which we obtain information on the shift of the centre of mass μ' of $A^{(k)}$ away from the target location and resolution lengths $w_{x',y',z'}$ in the three directions defined by the local Cartesian frame. Fitting is performed by using the Levenberg-Marquardt algorithm for solving the weighted least-squares problem

$$\underset{\mathbf{p}}{\operatorname{arg\,min}} \sum_{j} \mathcal{V}_{j} \left(A_{j}^{(k)} - \hat{g}^{(k)}(\mathbf{r}'_{j};\mathbf{p}) \right)^{2}, \qquad (9)$$

using the volume \mathcal{V}_j associated with each *j*-th grid node. For a practical initial guess, we use the specifications of the corresponding target kernel $T^{(k)}$ at the location of interest. As a remark, other parametrizations are possible that could approximate even more accurately the shape of averaging 596
 kernels. Fichtner & Trampert (2011) pointed out that Gram-597

⁵³⁸ Charlier expansions can be used to that end. These make it
⁵³⁹ possible to approximate functions, or more strictly distribu⁵⁴⁰ tions, like the averaging kernels from a series of their cumu⁵⁴¹ lants (e.g. mean, variance, third central moment and more
⁵⁴² complicated quantities at higher orders). While this can lead
⁵⁴³ to more accurate approximations of the complete function,

601 Fichtner & Trampert (2011) suggested that from a practical 544 602 point of view, these expansions might not always offer more 545 603 physically interpretable information than the plain Gaus-546 604 sian approximation itself. However, these and other suitable 547 605 parametrizations may be required for analysing functions 548 606 that are vastly different across the model domain, such as 549 607 kernels for the trade-off between different physical model pa-550 rameters (which is not the case for the $A^{(k)}$ with Gaussian 608 551 609 targets $T^{(k)}$ in this study). 552

610 The difference between the two Gaussian parametriza-553 611 tions $\hat{g}^{(k)}$ and $\tilde{g}^{(k)}$ is visualized schematically in Fig. 6. The 554 612 semi-axes of the ellipse at half the maximum are the es-555 613 timated resolution lengths $w_{x',z'}$ (+ $w_{y'}$ in the actual 3-D 556 614 ellipsoid), that can be compared to the target widths $w_{\rm H/V}$. 557 615 These estimates, as well as the remaining parameters in **p**, 558 616 may differ slightly between the two parametrizations. The 559 largest mean absolute deviation between $\hat{g}^{(k)}$ and $\tilde{g}^{(k)}$ we ⁶¹⁷ 560 618 found for any length estimate and given parameter combina-561 619 tion is ~ 16 km. However, in case of kernels with reasonably 562 good target fits these differences are small, and we noticed $\ ^{620}$ 563 that on global scales, the choice between uncorrelated and $^{\ 621}$ 564 correlated Gaussian has only a minor effect (see supporting 622 565 623 material). A remaining problem is the identification of com-566 plex averaging kernels that cannot be well approximated 567 by either of the Gaussian parametrizations. We therefore 625 568 propose a practical strategy to deal with this issue in the 569 following section. 570 626

571 4.2 Kernel classification based on the 'focus'

629 For a robust interpretation of the resolution informa-572 630 tion, one needs to know for every target location whether 573 631 the Gaussian-based estimates of resolution length actually 574 relate well to the shape and mass of the averaging kernel (i.e. the integral over $A^{(k)}$ for a given volume). Since the 632 575 576 centre of $\hat{g}^{(k)}$ is an estimate of the centre of mass in $A^{(k)}$, 634 577 the (main) peak regions of both functions should by design 578 overlap. If we can measure, by comparison of the masses 579 637 of the two kernels, if this is actually true for a given tar-580 get location, it would allow us to quickly identify for which 581 choices of the trade-off parameter η and target resolution 582 the resulting averaging kernels deviate significantly from a 583 638 Gaussian. 584

To get a diagnostic tool for this specific problem, we 639 585 develop in the following a global classification scheme for 640 586 the averaging kernels, measuring their quality with respect to the best-fitting Gaussian $\hat{g}^{(k)}$. First, it is necessary to 641 587 642 588 define the actual volume of the peak region of $\hat{g}^{(k)}$, for which 643 589 we can draw inspiration from the simpler 1-D case. For a 644 590 1-D Gaussian function, 76 per cent of the total area lies 645 591 within $\pm w$, the half width at half maximum (our measure for 646 592 resolution length), around the mean. In higher dimensions, 647 593 however, this well-known concept does not hold. In fact, the 648 594 integrated mass of $\hat{g}^{(k)}$ within the 3-D ellipsoid given by the 649 595

three half widths at half maximum $w_{x^\prime,y^\prime,z^\prime}$ (respectively the inner ellipse in Fig. 6) defined by

$$\left\{ \left(\frac{x' - \mu_{x'}}{w_{x'}}\right)^2 + \left(\frac{y' - \mu_{y'}}{w_{y'}}\right)^2 + \left(\frac{z' - \mu_{z'}}{w_{z'}}\right)^2 \le 1 \right\}$$
(10)

covers merely ≈ 29 per cent (if $w_{x'} = w_{y'} = w_{z'}$) of the total mass of the Gaussian. The exact value also depends on the specific half widths in each direction, and thus may vary for different $\hat{g}^{(k)}$. Since the SOLA kernels determine average values in volumetric regions, we consider 29 per cent of the total mass to be insufficient to properly describe the characteristics of $A^{(k)}$. Therefore, we aim to reproduce the 1-D convention, with roughly 76 per cent of the total mass, for the 3-D situation here as well. This can be achieved by scaling up the ellipsoid in eq. (10) and replacing the $w_{x',y',z'}$ with the corresponding half widths at one-eighth maximum from the best-fitting Gaussian $\hat{g}^{(k)}$ (the outer ellipses in Fig. 6). In practice, we determine the ellipsoid by evaluating $\hat{g}^{(k)}$ directly on the tomographic grid and including all nodes that hold values larger than one-eighth of the maximum (taken from the continuous function). The absolute mass $\hat{g}_{in}^{(k)}$ inside this spatial domain $E^{(k)}$ then approximately represents 76 per cent of the total mass $(\hat{g}_{in}^{(k)} + \hat{g}_{out}^{(k)})$. The exact number of the total mass $(\hat{g}_{in}^{(k)} + \hat{g}_{out}^{(k)})$. meric value, however, may still differ by a few per cent of the total mass depending on the size and location as well as on the error from evaluation on a discrete grid. Having formulated an expectation on the volume and mass $\hat{g}_{in}^{(k)}$ of the peak region, we can then separate the mass contribution of the associated averaging kernel $A^{(k)}$ within and outside the ellipsoid $E^{(k)}$:

$$\underbrace{\int_{\mathbf{r}' \in E^{(k)}} A^{(k)}(\mathbf{r}) \, d^3 \mathbf{r}}_{A_{in}^{(k)}} + \underbrace{\int_{\mathbf{r}' \notin E^{(k)}} A^{(k)}(\mathbf{r}) \, d^3 \mathbf{r}}_{A_{out}^{(k)}} = 1. \quad (11)$$

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Note again that the total mass of $A^{(k)}$ is equal to 1 in the SOLA method owing to the unimodular condition in eq. (4). Generally, both $A_{in}^{(k)}$ and $A_{out}^{(k)}$ include positive as well as negative kernel values. However, the negative contributions to $A_{in}^{(k)}$ were found to not exceed 2–3 per cent for any averaging kernel that we computed. For more than ~80 per cent of all averaging kernels, negative values in $A_{in}^{(k)}$ are ≤ 1 per cent and vanished completely for ~8-25 per cent (depending on the specific $T^{(k)}$ and η). We then define the 'focus' ξ of the averaging kernel based on the mass ratio of $A_{in}^{(k)}$ and $\hat{g}_{in}^{(k)}$ as

$$\xi = \left(\frac{A_{in}^{(k)}}{A_{in}^{(k)} + A_{out}^{(k)}}\right) / \left(\frac{\hat{g}_{in}^{(k)}}{\hat{g}_{in}^{(k)} + \hat{g}_{out}^{(k)}}\right) = \frac{A_{in}^{(k)}(\hat{g}_{in}^{(k)} + \hat{g}_{out}^{(k)})}{\hat{g}_{in}^{(k)}}.$$
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The normalization with the total mass $(\hat{g}_{in}^{(k)} + \hat{g}_{out}^{(k)})$ of $\hat{g}^{(k)}$ within the model domain allows one to take into account possible errors through discretization (for the absolute values of $\hat{g}_{in}^{(k)}$) and ellipsoids $E^{(k)}$ that are intersected by the surface. Using these relative mass contributions is helpful because it makes ξ a uniform metric for all $A^{(k)}$. In case that the unimodular condition for $A^{(k)}$ (eq. 4) is not fulfilled, e.g. when using other kernels that are not derived by a Backus–Gilbert type inversion, one should modify the focus ξ and normalize $A_{in}^{(k)}$ also with the total mass of the averaging kernel (i.e. one should use the more general ex-

pression of ξ in eq. 12). To finally classify the quality of the ros averaging kernels, we define 5 categories for ξ : right ros ros

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• $\xi < 0.5$, 'Not Focused':

 653 $A^{(k)}$ is hardly focused and likely degraded by multiple peaks 712 654 or strong side lobes, often due to either individual finite- 713 frequency kernels dominating the averages or strong sensi- 714 tivity in the upper mantle; 715

• $\xi \ge 0.75$, 'Sufficient (Suff.)':

Threshold between acceptable and unacceptable kernel fit 717 with respect to $\hat{g}^{(k)}$; ensuring that all sufficient $A_{in}^{(k)}$ are at 718 least more concentrated inside the Gaussian ellipsoid $E^{(k)}$ 719 rather than outside of it; 720

• $0.9 \le \xi < 1.1$, 'Good':

⁶⁶³ Deviation of $A_{in}^{(k)}$ from the ideal case is less than 10 per cent; ⁷²² ⁶⁶⁴ • $\xi = 1$, 'Ideal': ⁷²³

Relative mass contributions of $A^{(k)}$ and $\hat{g}^{(k)}$ within ellipsoid 724 $E^{(k)}$ are equal; 725

• $\xi \geq 1.1$, 'Highly Focused':

⁶⁶⁸ $A^{(k)}$ is more peaked, i.e. has significantly more mass within ⁶⁶⁹ the ellipsoid $E^{(k)}$ than $\hat{g}^{(k)}$.

This classification with the focus can be seen as a heuristic $_{\ensuremath{\text{726}}}$ 670 tool for deciding on the quality of the kernels, since a rela-671 tive redistribution of the mass of $A^{(k)}$ to outside the peak 672 728 region is clearly associated with a decrease in ξ . Therefore, 673 729 although these focus categories do not fully characterize the 674 730 detailed shape of a specific averaging kernel, they yet pro-675 731 vide a basic way to test whether the Gaussian approximation 676 732 is locally plausible. The classification scheme also gives an 677 733 indication for the fit to $T^{(k)}$, if $\hat{g}^{(k)}$ is not shifted or broadly 678 734 stretched beyond the target resolution length. This is be-679 cause averaging kernels that agree well with $\hat{g}^{(k)}$ tend to be 735 680 736 also centred. 681 737

4.3 Examples for resolution quantification and classification

Before we apply the previously introduced methods on a 684 742 global scale (see Section 4.4), we demonstrate more explic-743 685 itly how they act together to describe resolution. Therefore, 744 686 instead of averaging kernels that are obviously Gaussian-745 687 shaped (see Fig. 5), we consider two less-intuitive examples 746 688 in Fig. 7. Again, we employ a high-resolution and a low-747 689 resolution setup (Figs 7a and 7b, respectively). Unlike the 748 690 averaging kernels, the functions $\hat{g}^{(k)}$ are defined beyond the 749 691 mantle domain V, and the associated Gaussian ellipsoids 750 692 for the focus ξ (at max/8) may thus come close to or ex-751 693 tend through the surface. As a consequence, the best-fitting 694 752 Gaussian $\hat{g}^{(k)}$ in Fig. 7(a) can yield a vertical half width $w_{z'}$ 753 695 of 322 km that is considerably larger than the target length 754 696 of $w_{\rm V} = 200$ km. The estimated resolution lengths in hor-755 697 izontal direction of 370 km and 340 km (North-South and 756 698 West-East, respectively) are as well somewhat larger than 757 699 the actual target length of $w_{\rm H} = 300$ km. 758 700

The averaging kernel for the low-resolution case in Fig. 701 759 7(b) is even more complex. It exhibits localized smearing 760 702 of uppermost mantle structure from individual data ker-761 703 nels, and several distinct local maxima in and outside the $_{762}$ 704 target region. Resolution lengths estimated by $\hat{q}^{(k)}$ are in 763 705 horizontal direction 890 km and 864 km (North-South and 764 706 West-East) and in vertical direction 459 km. They are thus 765 707 horizontally narrower but vertically larger than the target 766 708

lengths of $w_{\rm H/V} = 900/400$ km. There is also a considerable shift of the centre of $\hat{g}^{(k)}$ from the target location of $\mu' = (-66, -97, -106)$ km, which might not be expected by merely inspecting the cross-sections at these azimuths. Overall, it is at least debatable whether the best-fitting Gaussian $\hat{q}^{(k)}$ is an adequate approximation in this case. For the $A^{(k)}$ in Fig. 7(a) we obtain a focus value $\xi = 0.89$ (i.e., certainly 'sufficient' and almost in the range of 'good' kernels), meaning that although the target shape is not fully matched, the best-fitting Gaussian $\hat{g}^{(k)}$ can be used with confidence for describing local resolution lengths. The averaging kernel in Fig. 7(b), however, is classified as insufficient, with a focus value $\xi = 0.62$ suggesting that maxima present outside the ellipsoid $E^{(k)}$ may have a significant impact on the corresponding model estimate for the average value. Further resolution estimates and focus values ξ for the averaging kernels in Fig. 3 can be found in the supplementary material.

4.4 Global classification of resolution

The results of our kernel classification on a global scale, for all previously used combinations of the trade-off parameter η and the target kernels $T^{(k)}$, are visualized in Figs 8 and 9. As expected, we find overall a larger number of Gaussian-shaped averaging kernels at smaller values for η . The percentage of acceptable $A^{(k)}$ ($\xi \ge 0.75$) is shown above each map, varying between 8-79 per cent at 595 km and 34-91 per cent at 1210 km depth. While there is a strict trend of fewer acceptable averaging kernels towards higher η , this is not necessarily the case when increasing the size of $T^{(k)}$. The largest number of acceptable kernels is actually obtained for the intermediate target kernel size with $w_{\rm H/V}$ = 600/300 km using η_1 (left column in Fig. 8). Consistently fewer acceptable kernels are found for the target resolution of $w_{\rm H/V} = 900/400$ km. Note that this difference between the target sizes does not imply that model averages are better constrained at the smaller scales rather than at large ones as we are not investigating the resolving power itself here, but the appropriateness of the Gaussian approximation. Instead, the classification maps suggest that one could lower the trade-off parameter η even further and obtain a larger number of Gaussian-shaped averaging kernels also for the larger target sizes. 'Highly focused' kernels (i.e. $\xi \geq 1.1$), are only observed for the HR inversion (Fig. 8, top left panel) at a few target locations east of Hawaii. This category suggests that narrower Gaussian estimates $\hat{g}^{(k)}$ may be possible in those cases and the associated resolution lengths could then be regarded as upper bounds of the size of the corresponding $A^{(k)}$. Alternatively, increased focus values could in some cases be explained by an overshooting local maximum at the peak combined with pronounced negative kernel values outside the peak region. The trend in Fig. 8 then also suggests that a larger value for η , with consequently lower model uncertainty, can be chosen locally if such a highly focused $A^{(k)}$ is not desirable (note again that the trade-off parameter can in principle be chosen for each target location individually). Given all the combinations of $T^{(k)}$ and n that we tested, a consistent classification of 'sufficient' and better is only possible in a few regions of excellent data coverage in the Northern Hemisphere.



Resolution in global seismic tomography 11

Figure 7. Averaging kernels $A^{(k)}$ and their best-fitting Gaussians $\hat{g}^{(k)}$ using (a) the high-resolution setup $(\eta_1, T^{(k)})$ with horizontal and vertical extent $w_{H/V} = 300/200$ km), below Western Australia and (b) the low-resolution setup $(\eta_3, T^{(k)})$ with $w_{H/V} = 900/400$ km), below Venezuela. Both target locations and the corresponding lateral cross-sections are at 595 km depth. Vertical cross-sections are marked by blue and yellow dots. Larger dots specify the negative direction for the local coordinate system in which resolution lengths and shifts of the centre are quantified. Additionally, the value for the focus ξ is shown, which serves as a metric for quantifying the appropriateness of $\hat{g}^{(k)}$ (see Section 4.2).

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767 5 GLOBAL RESOLUTION & UNCERTAINTY 768 MAPS

The tools presented in the previous section allow us to in-793 769 spect the resolution lengths for the varying tomographic pa-794 770 rameter setups, in connection with a basic test of the validity 795 771 of the Gaussian approximation. On a global scale, this has 796 772 the power to reveal concisely the impact of the different in-797 773 version parameters on resolution. In combination with the 798 774 propagated model uncertainties $\sigma_{\hat{m}^{(k)}}$, which are straightfor-775 wardly calculated with the SOLA method (see eq. 5), this 800 776 also makes it possible to uncover the locally varying trade- 801 777 offs between resolution and uncertainty. To illustrate this, we 802 778 show global maps of the estimated resolution lengths for all 803 779 inversion parameter combinations in Figs 10–13. We apply 804 780 our classification scheme to mask all 'insufficiently' focused 781 805 averaging kernels in these maps. In addition, the mean and 806 782 standard deviation of all resolution lengths, given by averag-783 807 ing kernels classified as 'sufficient' and better, are specified 808 784 above each map. 809 785

For vertical resolution estimates at 595 km depth, one can observe a strong variability for the case with target lengths of $w_{\rm H/V} = 300/200$ km (Fig. 10, top row). At this target size, the mean vertical resolution length ranges be-

tween 316-329 km across all trade-off values and with increasing values of η , one can see a clear progression towards larger vertical extent of the averaging kernels. Minimummaximum values of vertical resolution for each target size $w_{\rm V} = 200, 300, 400 \text{ km}$ are 211–501 km, 321–521 km and 416–534 km, respectively. The maps therefore show that the vertical target length only is approached in regions of high data coverage, but overall cannot be reached by any 'sufficient' averaging kernel. Numerous $A^{(k)}$ deviate strongly (>100 km) from the vertical target resolution in Fig. 10, but they do not all necessarily fall in the category of 'insufficient' averaging kernels (e.g. beneath the North Pacific). This means that a useful Gaussian $\hat{g}^{(k)}$ was obtained, although a more accurate fit to the target kernel $T^{(k)}$ could not be achieved at these locations. We chose to additionally indicate the non-Gaussian $A^{(k)}$ as shaded areas to roughly analyse the range of estimated lengths there, even if they are less reliable. The vertical resolution lengths in those regions are in fact often in line with surrounding acceptable $A^{(k)}$, but may as well be anomalously low or high, quite strikingly for instance around the East Pacific Rise (we analyse an example for this region in more detail in Section 6.1). An opposite trend can be observed in the mid-mantle at 1210 km depth in Fig. 11, where the prescribed target lengths are overall



Figure 8. Classification of averaging kernels based on the focus ξ at a depth of 595 km depth for different target resolution lengths and trade-off parameters η . Blue colours mark averaging kernels that are sufficiently ('Suff.') well approximated by the best-fitting Gaussian $\hat{g}^{(k)}$, meaning that resolution length can be robustly extracted from $\hat{g}^{(k)}$. 'Good' averaging kernels are subdivided in at levels of $\xi = 0.9, 0.95, 1.0, 1.05$ to reveal their deviation from the ideal case in more detail. Red colours accordingly indicate more complex averaging kernels $A^{(k)}$ that are often affected by strong kernel side lobes. Percentages above the maps display the relative amount of acceptable $A^{(k)}$.



Figure 9. Classification of averaging kernels based on the focus ξ at a depth of 1210 km.



Resolution in global seismic tomography 13

Figure 10. Vertical resolution lengths of the averaging kernels at 595 km depth, as estimated from the best-fitting Gaussian $\hat{g}^{(k)}$. Shaded areas mark the regions where the classification from Fig. 8 indicates that the $A^{(k)}$ are insufficiently Gaussian-shaped and that resolution length must be interpreted with caution there. Mean values and standard deviations of the resolution lengths of all 'sufficient' averaging kernels are given above each map. Contour lines are drawn at 200, 300, 400 and 500 km.



Figure 11. Vertical resolution lengths of the averaging kernels at 1210 km depth. Same as Fig. 10 with shaded areas given by Fig. 9.



Figure 12. Longitudinal resolution lengths of the averaging kernels at 595 km depth, as estimated from the best-fitting Gaussian $\hat{g}^{(k)}$. Same layout as Fig. 10.



Figure 13. Longitudinal resolution lengths of the averaging kernels at 1210 km depth. Same layout as Fig. 11.

fitted well on a global scale. At this depth, the mean values fitted well on a global scale. At this depth, the mean values for the estimated vertical resolution lengths are however of estimates in the smaller compared to their target, and the full range of for estimates is 176–239 km, 261–336 km and 351–423 km for for were estimates is 176–239 km, respectively. The state of the

Maps of longitudinal resolution lengths for the same 819 877 depths are shown in Figs 12 and 13. The global mean values 878 820 for each selection of $w_{\rm H/V}$ and η as well as the associated 879 821 distributions of longitudinal resolution lengths are quite sim-880 822 ilar at both depths. The respective target lengths are well 881 823 approached in many regions. At 595 km depth, minimum-824 maximum values of longitudinal resolution for the different 883 825 target resolutions $w_{\rm H} = 300, 600, 900$ km (again across all 884 826 values of η and for all 'sufficient' and better averaging ker-827 nels) are 253-673 km, 495-826 km and 778-1082 km, respec-886 828 tively. At 1210 km depth, a larger number of 'sufficient' aver-829 887 aging kernels relative to the total number per layer is found. 888 830 Still, the corresponding ranges are slightly narrower with 889 831 minimum-maximum values of 293-581 km, 560-778 km and 890 832 828–1104 km. Fairly similar maps and trends can be found 891 833 for the estimated latitudinal resolution lengths (shown in 892 834 Appendix C, Figs C1 and C2). 803 835

To complete the results for our tests of SOLA inversion⁸⁹⁴ 836 parameters, we show the associated global maps of prop-895 837 agated model uncertainty $\sigma_{\hat{m}^{(k)}}$ in Figs 14 and 15. As in-896 838 tuitively expected, the uncertainty increases systematically 897 839 with lower values of η , but also with smaller target sizes. A 898 840 841 possible reason for this could be that fewer finite-frequency 899 kernels interact within the smaller target kernel volume in $_{\rm 900}$ 842 this case. The highest global mean and largest variability 843 901 thus correspond to the HR inversion setup ($\eta_1, w_{\rm H/V} =$ 844 902 300/200 km) with $\sigma_{\hat{m}^{(k)}} = 0.668 \pm 0.138 \ \%$ at 595 km depth, 903 845 and 0.478 ± 0.112 % at 1210 km depth. For the largest target 846 904 kernels with $w_{\rm H/V} = 900/400$ km, the model uncertainties 905 847 are vanishingly small. This indicates that lower values for 906 848 η should be chosen at larger target scales compared to the 907 849 best choice of η at smaller target resolution. Overall, there 908 850 is no strict pattern emerging for the variability of the model 851 909 uncertainty with respect to the associated resolution length 910 852 estimates. Instead, one can observe regions of reduced $\sigma_{\hat{m}^{(k)}}$ 911 853 both where the potential for overall high resolution is ob-912 854 served (e.g. around Japan, with 'ideal' focus), but also where 913 855 our classification tool suggested a poor Gaussian approxima-856 914 tion of the averaging kernels. 857 915

858 6 DISCUSSION

859 6.1 Role of the Gaussian approximation

As shown in Section 4, an approximation of the averaging 920 860 kernels $A^{(k)}$ by some parametric function is crucial for deter-921 861 mining their resolving power shown in our global resolution 922 862 maps. In this respect, Gaussian functions are a convenient 923 863 choice and serve as a tool for both estimating resolution 924 864 lengths and for identifying kernels of more complex shape 925 865 and consequently low focus. Earlier studies have also relied 926 866 on the Gaussian approximation in order to parametrize the 927 867 resolution matrix (An 2012) or the point-spread functions 928 868 based on the Hessian matrix (Fichtner & Trampert 2011) 929 869 prior to the actual inversion step. In contrast to this earlier 930 870 work, we computed entire sets of averaging kernels without 931 871

additional assumptions on their structure, and tested subsequently if a reduction of their complexity in the form of a Gaussian parametrization is warranted. Although using the SOLA method and targeting Gaussian-shaped averaging kernels here, it was not possible to obtain robust resolution proxies from the kernels at every location and at reasonable levels of uncertainty (highlighted by the classification in Figs 8, 9 and uncertainty maps in Figs 14, 15). We therefore suggest that our concept should be applied to new datasets, including hypothetical ones, also to investigate potential improvements in resolution resulting from additional seismic stations at new locations.

The reliability of our resolution length analysis, including the computation of the kernel focus ξ , clearly depends on the quality of the individual Gaussian approximations. For example, an irregular feature was discovered by the classification scheme at a depth of 595 km (Fig. 8), suggesting strongly non-Gaussian averaging kernels around the East-Pacific Rise. Across this region, the kernels are insufficiently focused and show notably low estimates of vertical resolution lengths for all inversion parameter combinations (Fig. 10). In Fig. 16, we provide an example of such a kernel using η_1 and target lengths $w_{\rm H/V} = 300/200$ km (i.e. the 'highresolution' scenario). It is strongly shifted to greater depths with respect to the target location by about 190 km, as estimated from the centre of the best-fitting Gaussian. While the Gaussian appears to fit this kernel well at the level of the half width at half maximum, the low value for the focus of $\xi = 0.25$, however, suggests a poor approximation. This is mainly due to elevated averaging values and oscillations of the kernel outside the plane of the cross-section (indicated for example by the local maximum visible in the lateral cross-section in Fig. 16). The estimated resolution lengths may thus still be reasonable even for these complex kernels, while the focus successfully points out their inadequacy for describing comprehensively the resolution characteristics. Again, this demonstrates why the computation of resolution lengths from 2-D cross-sections can be potentially misleading. A more robust assessment of resolution lengths for these complex scenarios instead requires some combination of accurate estimates, based for example on the 3-D Gaussian parametrization, and a classification that we realize using the concept of 'focus' here. Our analysis illustrates that both together are a useful way to extract the relevant information on resolution reliably from a large set of averaging kernels.

6.2 Implications for SOLA tomography

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A key result of Zaroli (2016) was that choosing a constant value for the trade-off parameter η per layer can produce coherent SOLA tomographic images (i.e. showing geodynamically interpretable large-scale features) with bounded uncertainty on a global scale. In contrast to Zaroli (2016), in which spheroidal, constant target functions of variable size adapted to an irregular data-driven grid were used, we chose a laterally (almost) equidistant model parametrization and tested a homogeneous Gaussian target kernel size consistently at all model locations. Our results show that, for a given selection of inversion parameters η and $T^{(k)}$, the model uncertainties $\sigma_{\hat{m}^{(k)}}$ from propagated data errors may have a low variability across all target locations. This sup-



Figure 14. Propagated model uncertainty $\sigma_{\hat{m}^{(k)}}$ at 595 km depth for different target kernel sizes and trade-off parameters. Dashed contour lines are only drawn at the values shown in the colour bar. Values above each map show the global mean \pm standard deviation.



Figure 15. Same as Fig. 14 (model uncertainty $\sigma_{\hat{m}^{(k)}})$ at a depth of 1210 km.



Figure 16. Example of a poorly centred averaging kernel $A^{(k)}$ 981 in the vicinity of the East-Pacific Rise for the high-resolution 982 inversion setup (η_1 , target resolution $w_{\rm H}/w_{\rm V}$ = 300/200 km). 983 The lateral cross-section at 595 km (top plot) does not reveal 984 the main peak due to its vertical shift of about -190 km and the 985 limited vertical half width of 105 km. The kernel is insufficiently 986 focused ($\xi < 0.75$), which implies that its best-fitting Gaussian 987 is not representative for the complete averaging volume. Ellipses 988 and contour lines are explained in Fig. 7. 989

ports the notion that, from the perspective of error propa-991 932 gation, it is indeed viable to use one single trade-off parame-992 933 ter (e.g., here η_2 with the current dataset) per tomographic 993 934 layer. In previous applications of the SOLA method it has so 994 935 far not been clear, however, whether the averaging kernels 995 936 also reasonably fit the target kernels $T^{(k)}$ at a given level 937 996 of uncertainty. The classification scheme developed here of-938 997 fers additional guidance in this regard by making it possible 998 939 to assess the quality of the averaging kernels with respect $_{\ 999}$ 940 to a best-fitting 3-D Gaussian instead of the target kernel. 1000 941 For certain research questions it might be further on de- 1001 942 sirable to obtain global tomographic models with different 1002 943 characteristics, favouring either high-resolution, uniformly 1003 944 good 'focusing' of averaging or resolving kernels, low uncer- 1004 945 tainty $\sigma_{\hat{m}^{(k)}}$, or a regionally variable mix thereof. It is, for 1005 946 example, not clear whether geodynamic inversions that aim 1006 947 at retrodicting past mantle evolution would benefit from ei- 1007 948 ther homogeneous tomographic resolution or from homoge-1008 949 neous model uncertainty (see e.g. Colli et al. 2020). This in 1009 950 turn requires the joint adaptation of target kernels $T^{(k)}$ and 1010 951 trade-off values $\eta^{(k)}$ for every individual parameter location. 1011 952 Such an analysis is beyond the scope of the current study, 1012 953 but can readily be tackled with the tools developed here. 1013 954 For the given shape of the target kernels in terms of a Gaus- 1014 955 sian, we have demonstrated that it becomes clear from just 1015 956 a few inversions, which range of values for $\eta^{(k)}$ is practically 1016 957 relevant and whether another size or shape for the target 1017 958 function $T^{(k)}$ should be employed. This empirical procedure ${\scriptstyle 1018}$ 959 is necessary because no automatic criterion or rule exists to 1019 960 determine the 'ideal' SOLA inversion parameters at every 1020 961 target location without excessive testing. For global appli-1021 962 cations, this would be computationally prohibitive. The joint 1022 963 analysis of model uncertainties $\sigma_{\hat{m}^{(k)}}$ and the focus ξ , which 1023 964

we used to set up the classification scheme for the averaging kernels, in comparison, is practically viable and provides essential information for selecting useful local combinations of $T^{(k)}$ and $\eta^{(k)}$, globally.

969 7 CONCLUSIONS

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We have presented a practical concept and its application for the assessment of resolution and uncertainty of tomographic images on a global scale. It is based on: 1) explicitly available averaging kernels and uncertainties, here enabled by the SOLA method, 2) a 3-D Gaussian parametrization of the averaging kernels for estimating resolution lengths consistently, and 3) a classification scheme for identifying regions where the Gaussian approximation may not accurately represent the spatial averaging. Through this combination, it is possible not only to investigate and visualize the resolution information for all the averaging kernels together, but also to indicate in a straightforward fashion in which regions the results can be reliably interpreted. At the same time, it shows in which locations specific care must be taken, for example where individual analysis of the local resolving power of the given dataset may be advisable.

Instead of interpreting model values, we employed the approach presented here for testing various combinations of SOLA inversion parameters in terms of their effects on final resolution and uncertainty, as there is no formal way to determine any 'ideal' setup. Our analysis shows that only a few inversions are required for a given realization of the target kernels to pinpoint whether their size or shape needs to be adapted and which range of values for the trade-off parameter is useful. This in turn is important for effectively selecting the proper combinations of these parameters at each target location in case of specific requirements (e.g. tuning towards more homogeneous resolution or more homogeneous model uncertainty). Finally, we emphasize once more that the general analysis performed here as well as the classification scheme are not restricted to the setup based on SOLA. They can also be applied to the closely-related resolution concept for point-spread functions, in case they are explicitly available.

In summary, the analysis with our tomographic framework revealed conclusively that resolution lengths from the SOLA method can be primarily tuned by the choice of target kernel size, and only to a lesser extent by the trade-off parameter. Despite this design control, a good focusing of the averaging kernels (i.e. adequacy of the approximation with a Gaussian) cannot be guaranteed on a global scale with the data and possible inversion setups employed here (especially for a target resolution going down to ~ 300 km horizontally and ~ 200 km vertically). Most notably, as expected with body waves, one has less control on vertical than on horizontal resolution length, especially at shallower depths in the mantle. However, a laterally homogeneous distribution of uncertainty is generally possible by choosing locally varying trade-off parameters. Additionally, the SOLA uncertainties remind us that even if averaging kernels or point-spread functions with high resolution could be obtained everywhere with specific regularization choices, the estimated model may be highly speculative in regions of poor data coverage. It is for these reasons that a proper and

coherent quantification of resolution and model uncertainty 1079
 is of critical importance, since this is a prerequisite to better 1080
 inform independent geophysical studies that rely on global 1081
 tomographic images.

1028 ACKNOWLEDGMENTS

We thank the editor Carl Tape, as well as Nathan Simmons ¹⁰⁸⁸ 1029 and an anonymous reviewer for their constructive comments 1030 1090 that improved the clarity of our manuscript. This work was 1031 supported by the Deutsche Forschungsgemeinschaft (DFG) 1032 1092 under grant SCHU 2914/7-1 (project ID 456788150). We are $\frac{1}{1093}$ 1033 grateful to the High-Performance Computing Centre of the 1094 1034 University of Strasbourg and the Leibniz-Rechenzentrum 1095 1035 (LRZ) for providing support and access to computing re- 1096 1036 sources. We also want to thank Jens Oeser for an excellent 1097 1037 in-house computational infrastructure at LMU Munich. This 1098 1038 work was partly supported by the Programme National de ¹⁰⁹⁹ 1039 1100 Planétologie (PNP) of CNRS/INSU, co-funded by CNES. 1040 1101

1041 DATA AVAILABILITY

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1042The data underlying this article will be shared on reasonable11061043request to the corresponding author.1107110811081108

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APPENDIX A: USING BACKUS-GILBERT THEORY FOR TOMOGRAPHY

In the original theory, Backus and Gilbert propose to construct optimal kernels $A^{(k)}(\mathbf{r})$ by approximation of a delta peak $\delta(\mathbf{r} - \mathbf{r}^{(k)})$ in terms of a specific deltaness criterion (Backus & Gilbert 1968), for example by minimizing the socalled spread (Backus & Gilbert 1970). For a perfect delta peak, the linear averaging in eq. (1) would simply show that $\hat{m}^{(k)} \approx m(\mathbf{r}^{(k)})$, which is however only possible if the data were complete and free of errors. Instead, one will typically need to accept $A^{(k)}(\mathbf{r})$ that deviate from $\delta(\mathbf{r} - \mathbf{r}^{(k)})$, due to the available set of data kernels $K_i(\mathbf{r})$, and in order to moderate the propagation of data errors into the inferred averages. A tomographic model would then consist in a collection of M local averages, $\hat{m}^{(k)}$, for k = 1, ..., M. Rewriting eq. (1) in a discrete notation with volumetric weights \mathcal{V}_j appropriate for each j-th grid node as

$$\hat{m}^{(k)} = \sum_{j} A_j^{(k)} \mathcal{V}_j m_j = \sum_{j} \mathcal{R}_{kj} m_j , \qquad (A.1)$$

one can see that the averaging kernels projected onto a discrete model parametrization represent a single row of the resolution operator \mathcal{R}_k . The operator \mathcal{R} can then be used to retrieve information on the resolution for a specific model parametrization, while $A^{(k)}(\mathbf{r})$ essentially refers to the local resolution for the continuous model $m(\mathbf{r}^{(k)})$ (Trampert 1998). Theoretically, however there is no need to discretise the model, as the Backus–Gilbert approach essentially solves a continuous inverse problem. This can be exploited fully for example by the 'parameter-free' SOLA approach (Zaroli 2019), while in the present paper we still make use of the discrete formulation as described in Zaroli (2016). A noteworthy limitation is that the theory does not guarantee that the collection of averages together actually explains the data. We tested this for our class of models that use only body waves at the moment, and we observed that the global misfit reduction can actually be comparable to and sometimes even be better than in classic damped-least squares inversion with model norm damping. However, the question of data misfits in SOLA tomographies, especially for different inversion parameter choices, might be a relevant subject for future investigations.

APPENDIX B: EFFICIENTLY SOLVING THE SOLA B-G SYSTEM

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1466 Zaroli (2016) explained how (the discrete version of) the 1407 1467 SOLA optimization problem in eq. (3) can be solved together 1408 1468 with the unimodular constraint eq. (4) using a least-squares 1409 1469 approach. Therefore, the SOLA system can be rewritten in 1410 1470 the following fashion (see Appendix A1, Zaroli 2016): 1411 1471

$$\hat{\mathbf{x}}^{(k,\eta)} = \underset{\hat{\mathbf{x}}^{(k)}}{\arg\min} \left\| \begin{bmatrix} \mathbf{Q}^{(\eta)} \\ \eta \mathbf{I}_{N-1} \end{bmatrix} \hat{\mathbf{x}}^{(k)} - \begin{bmatrix} \mathbf{y}^{(k,\eta)} \\ \mathbf{0}_{N-1} \end{bmatrix} \right\|_{2}^{2} \cdot (B.1)^{\frac{1472}{1473}}$$

Both the matrix $\mathbf{Q}^{(\eta)}$, of size $(M+1) \times (N-1)$, and the right-1413 hand side $\mathbf{y}^{(k,\eta)}$, which incorporates the target kernel, de- 1474 1414 pend on the choice of a particular trade-off parameter $\eta.$ $_{^{1475}}$ 1415 The complete solution $\mathbf{x}^{(k)}$ can finally be recovered from 1416 1476 $\mathbf{\hat{x}}^{(k)}$ (the intermediate solution in eq. B.1). Considering our 1417 tomographic grid and dataset, in each inversion we need to $^{\rm 1477}$ 1418 solve for the Backus–Gilbert coefficients $\mathbf{x}^{(k)}$ with a SOLA 1478 1419 system matrix $\mathbf{Q}^{(\eta)}$ of size 153,324 \times 79,764 (~2 per cent ¹⁴⁷⁹ 1420 non-zero elements, ${\sim}2$ GB). As one needs to perform a single 1480 1421 1481 inversion for every grid node, computational costs for empir-1422 1482 ically testing inversion parameters can thus quickly become 1423 prohibitive. The SOLA method has the computational ad-1424 vantage that the left-hand side of the corresponding linear 1425 system in eq. (B.1) is independent of the target location $\mathbf{r}^{(k)}$, 1426 i.e. for a given η the SOLA system matrix $\mathbf{Q}^{(\eta)}$ does not 1427 change. This enables perfectly parallel computation over all 1428 M model parameters. To this end, one could simply increase 1429 the number of processors P at the cost of having to store $2 \cdot P$ 1430 times the SOLA system matrix **Q**. This obviously becomes 1431 problematic if the available computing system is limited in 1432 RAM, especially if the tomographic systems become even 1433 larger than the ones considered in this study. Alternatively, 1434 using improved parallel solvers based on LSQR (Huang et al. 1435 1436 2013; Lee et al. 2013) or using instead efficient direct methods (Bogiatzis et al. 2016) are other possible options that 1437 we considered and list here for documentation. We decided 1438 to solve eq. (B.1) by using a GPU version of the algorithm 1439 LSMR (Fong & Saunders 2011). We use LSMR as imple-1440 mented by Krylov.jl (a package of selected Krylov meth-1441 ods written in the programming language Julia, see Mon-1442 toison & Orban 2023), where, conveniently, no significant 1443 code changes are required compared to the CPU version. 1444 Once the GPU compute kernel is compiled, different left-1445 hand sides $\mathbf{y}^{(k,\eta)}$ can be asynchronously copied from a CPU 1446 to the GPU and solution vectors can be efficiently recovered 1447 for each run without additional time spent for data trans-1448 fer or solver setup. This way, performing the inversion for a 1449 single model parameter in our computations was about 50-1450 100 times faster on the GPU (using double precision) com-1451 pared to a single CPU. Also, this only requires that $\mathbf{Q}^{(\eta)}$, 1452 a large but highly sparse matrix, fits twice into GPU mem-1453 ory. As a remark, LSMR is recommended to be used over 1454 LSQR by Fong & Saunders (2011) if iterations have to be 1455 stopped early. This could theoretically be exploited to keep 1456 the solver time limited in case a reasonable maximum num-1457 ber of iterations for all inversions globally is found. However, 1458 we have not drawn on this here and relied on the stopping 1459 criteria suggested by Fong & Saunders (2011). For a given 1460 trade-off parameter η , we found that the time to solution 1461 was practically constant for all model parameters and over-1462 all increased roughly by a factor of 2 for an equal decrease 1463

in η (i.e. solutions for $\eta_1 = 5$ took about twice as long as for $\eta_2 = 10$). The use of GPUs therefore proves to be very useful for SOLA inversions with a least-squares approach and can greatly reduce the time required for computing a complete model with M parameters. As a final note, we used 2 NVidia RTX A5000 in this study, but modern GPU clusters and supercomputers often feature many more units. From a computational point of view, it should thus be straightforward to use larger grids and datasets with the SOLA method than presented here.

APPENDIX C: ADDITIONAL GLOBAL RESOLUTION MAPS

To complete the global analysis, we show additionally the estimated latitudinal resolution lengths in Figs C1 and C2. Plots for the horizontal and vertical shifts of the centres of mass of the averaging kernels can be found in the supplementary (online) material, as well as comparisons between the uncorrelated and correlated (i.e. including rotation) 3-D Gaussian estimates.



Figure C1. Latitudinal resolution lengths of the averaging kernels at 595 km depth, as estimated from the best-fitting Gaussian $\hat{g}^{(k)}$. Same layout as Fig. 10. As shown in the main text, shaded areas highlight regions of low 'focus' (i.e. the Gaussian is inadequate to reliably estimate resolution length for the averaging kernel).



Figure C2. Latitudinal resolution lengths of the averaging kernels at 1210 km depth. Same layout as Fig. 11.